

Kinetic mechanisms for the growth of drift-ballooning modes

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Kinetic mechanisms for the growth of drift-ballooning modes are studied theoretically. It is found that these modes may be unstable as a result of their resonant and collisional interactions with ions. The instability growth rates depend on the relation between the temperature gradient and the density gradient.

INTRODUCTION

It follows from single-fluid MHD with an oblique (magnetic) viscosity that as the toroidal wave number is increased unstable ideal ballooning modes become stable drift-ballooning modes, because of a stabilizing effect of the finite ion Larmor radius (Ref. 1, for example). From the standpoint of kinetic theory, the meaning is that only the hydrodynamic stability of the drift-ballooning modes is involved here. In this connection it is important to determine whether a kinetic growth of these modes is possible by virtue of their resonant or collisional interaction with plasma particles. That question is the subject of the present paper.

Since the drift-ballooning modes are associated with Alfvén wave branches, the instabilities with which we are concerned here may be thought of as certain versions of the class of kinetic Alfvén instabilities of a tokamak plasma. The study of the stabilities of this class was begun in Ref. 2, where an analysis was made of the resonant interaction of local Alfvén waves modified by effects of the magnetic-field curvature, with untrapped ions. The analysis of Ref. 2 dealt with the case of perturbations with $\epsilon^{1/2}v_{Ti}/qR < \omega < v_{Ti}/qR$, where ω is the wave frequency, $\epsilon = a/R$ is the inverse aspect ratio, a and R are respectively the minor and major radii of the torus, $v_{Ti} = \sqrt{2T_i/M_i}$ is the ion thermal velocity, T_i , and M_i are respectively the temperature and mass of an ion, and q is the safety factor. The condition $\omega > \epsilon^{1/2}v_{Ti}/qR$ allows us to ignore the weakening of the resonant interaction which results from the trapping of particles in regions of weak magnetic field (on the outer side of the torus). In Ref. 3, in contrast with Ref. 2, a study was made of the resonant interaction of hydromagnetic perturbations stabilized by the effects of the finite ion Larmor radius, with untrapped ions, in the opposite limit, $\omega < \epsilon^{1/2}v_{Ti}/qR$, in which this trapping of particles is extremely important. Also of interest for the problem at hand is the collisional interaction of waves with untrapped ions under the condition $\omega < v_i/\epsilon$, where v_i is the rate of ion col-

lisions. According to Ref. 2, this process can be described in terms of a longitudinal ion viscosity calculated in the banana regime. The role played by the process in the ideal $m = 1$ kink modes (the index of the azimuthal harmonic) was studied in Ref. 2. Waves with $\omega > v_i/\epsilon$ were also studied there.

In Sec. 2 we present the basic equation and derive a general dispersion relation for drift-ballooning modes, incorporating the effects listed above. This relation is analyzed in Sec. 3. Conclusions are drawn in Sec. 4.

2. INITIAL EQUATIONS

We start from the equation for ideal ballooning modes, with drift effects and effects of the interaction with ions (cf. Ref. 4, for example):

$$\frac{d}{dt} \left(t^2 \frac{d\xi}{dt} \right) - U_0 \xi + t^2 \frac{\omega^2}{\omega_A^2} \left(1 - \frac{\omega_{pi}^2}{\omega} \right) + \sigma = 0.$$

Here $\omega_A = Sc_A/qR$ is the Alfvén frequency, $q'a/q$ is the shear, $U_0 = 4\epsilon^2\beta_p \times (1 - 1/q^2 + 3\epsilon^2\beta_p/2)/S$ is the magnetic well, β_p is the ratio of the pressure of the poloidal magnetic field, ω_{pi} is the ion drift frequency in terms of the pressure gradient, $t = Sy$, and y is the standard ballooning variable. The function ξ is a Fourier component of the radial displacement of the plasma. The term σ incorporates effects of the resonant and collisional interactions of ions with waves and is defined by

$$\sigma = - \frac{2t}{\rho_0 R \omega_A^2} (\delta \tilde{p}_k \sin \theta)^{(0)},$$

where ρ_0 is the plasma density, $\delta \tilde{p}_k$ is a Fourier component of the oscillatory part of the perturbed

pressure, and $(\dots)^{(0)} = \int_0^{2\pi} (\dots) d\theta/2\pi$.

We write the perturbed pressure in the

$$\delta\bar{p} = \frac{M_i}{4} \int v^2 f dv, \quad (3)$$

where $\tilde{f} = \tilde{f} \exp(im\theta)$ is the perturbed part of the distribution function, and θ is the poloidal angle. To find the function \tilde{f} we use the drift kinetic equation (Ref. 2, for example)

$$-i\omega\tilde{f} + \frac{v_{\parallel}}{qR} \frac{\partial \tilde{f}}{\partial \theta} + i\omega \frac{M_i}{2T} v^2 \left(1 - \frac{\hat{\omega}'}{\omega}\right) \frac{a}{mR} F \sin \theta \tilde{f}' = \text{St}(\tilde{f}). \quad (4)$$

Here v_{\parallel} is the velocity of the particles along the magnetic field, $\hat{\omega}' = \omega_{pi} [1 + \eta(M_i v^2 / 2T)^{1/2}]$, ω_{pi} is the drift frequency in terms of the pressure gradient, $\eta = \partial \ln T / \partial \ln n_0$, F is the equilibrium Maxwellian distribution function (normalized to the particle density n_0 , as usual), St is the collision term, and the prime denotes a derivative with respect to the minor radius.

Introducing a function h , by analogy with Ref. 2,

$$\tilde{f} = -ihF \left(1 - \frac{\hat{\omega}'}{\omega}\right) \frac{a}{mR} \hat{h}', \quad (5)$$

and using (4) and (5), we find an equation for h :

$$-i\omega h + \frac{v_{\parallel}}{qR} \frac{\partial h}{\partial \theta} - i\omega \frac{M_i}{2T} v^2 \sin \theta = \text{St}(h). \quad (6)$$

Transforming to the ballooning representation, and using (2), (3), and (5), we then find

$$\sigma = \epsilon^2 \frac{\omega^2}{\omega_A^2} \Lambda \hat{h}, \quad (7)$$

where (cf. Ref. 2)

$$\Lambda = -\frac{1}{2\omega^2 R^2 q^2 n_0} \left(\int v^2 F h dv \sin \theta \right)^{(*)}. \quad (8)$$

We see that the effects of the collisional and resonant interaction of ions with drift-ballooning modes lead to a renormalization of the inertial term in Eq. (1) (cf. Refs. 2 and 5). Making use of this circumstance, and also using the procedure of joining the asymptotic solutions of Eq. (1) with those of Eq. (1) in the "inertialess" region, we find the dispersion relation [cf. Eq. (26) of Ref. 6]

$$\omega^2 (1 - \omega_{pi} / \omega + \Lambda) = -\gamma_0^2, \quad (9)$$

where γ_0 is the growth rate of the drift-ballooning modes, given explicitly in Ref. 4.

The problem has now been reduced to one of calculating the quantity Λ in the various regimes from (6) and (8) and solving the dispersion relation (9).

3. SOLUTION OF THE DISPERSION

3.1. Resonant interaction. In accordance with the Introduction, we consider two cases of the resonant interaction of wave with untrapped ions. The first case deals with the region of perturbation frequencies² $\epsilon^{1/2} v_{Ti} / qR < \omega < v_{Ti} / qR$; the second deals with perturbations with frequencies³ $\omega < \epsilon^{1/2} v_{Ti} / qR$.

3.1.1. The case of frequencies $\epsilon^{1/2} v_{Ti} / qR < \omega < v_{Ti} / qR$. We solve kinetic Equation (6) as in Ref. 2.

Incorporating drift effects, we find an ϵ for Λ :

$$\Lambda = i \frac{\sqrt{\pi} q v_{Ti}}{2 R \omega} \left[1 - \frac{\omega_{pi}}{\omega} \left(1 + \frac{3}{2} \eta \right) \right].$$

Let us consider the very simple case γ_0 . This condition means that the MHD instability is suppressed by the effects of the finite ion η us. From (9) and (10) we then find

$$\omega = \omega_{pi} \left(1 + \frac{3}{2} \eta \right) + i \sqrt{\pi} \frac{\omega_{pi}^2 R}{q v_{Ti}} \left(1 + \frac{3}{2} \eta \right) \eta.$$

An instability occurs if

$$\eta > 0 \text{ and } \eta < -1/3.$$

The ratio of the imaginary and real part frequency is $\sim \omega_{pi}^2 R / q v_{Ti}$.

3.1.2. The case of frequencies $\omega < \epsilon^{1/2} v_{Ti} / qR$

In this case, making use of results from Ref. 2, we find the following expression for Λ :

$$\Lambda = \frac{3q^2}{8\sqrt{2}\epsilon} \lambda,$$

where

$$\lambda = 1 - \frac{\omega_{pi}}{\omega} + i \frac{\sqrt{\pi}}{18\sqrt{2}} \left(\frac{\omega}{\omega_b} \right)^2 \left[1 - \frac{\omega_{pi}}{\omega} \left(1 - \frac{3}{2} \eta \right) \right]$$

By analogy with Ref. 3, we find two waves in the case $\gamma_0 \ll \omega_{pi}$:

$$\omega_1 = \omega_{pi} - i \frac{\eta}{1 + \eta} \frac{5\sqrt{\pi}}{36\sqrt{2}(1 + 8\sqrt{2}\epsilon/3q^2)} \left(\frac{\omega_{pi}}{\omega_b} \right)^2 \omega_{pi},$$

$$\omega_2 = \omega_{pi} + i \frac{3\eta - 2}{1 + \eta} \frac{\sqrt{\pi}}{36\sqrt{2}(1 + 8\sqrt{2}\epsilon/3q^2)} \left(\frac{\omega_{pi}}{\omega_b} \right)^2 \omega_{pi},$$

where

$$\omega_{pi}^2 = \frac{\gamma_0^2}{\omega_{pi} (1 + 3q^2/8\sqrt{2}\epsilon)}.$$

The first branch is unstable if

$$-1 < \eta < 0,$$

and the second if

$$\eta > 1/3 \text{ and } \eta < -1.$$

The ratio of the growth rate to the real part of the frequency is $\sim (\omega_{pi} / \omega_b)^2$.

Expression (15) for the frequency ω differs from that given in Ref. 3 but refers to no Mercier perturbations but also ballooning perturbations. In the case in which γ_0 is the rate of Mercier perturbations, expression (15) reduces to that derived in Ref. 3. The ballooning mode growth rate, can also be substituted as γ_0 .

3.2. Collisional interaction. Again in this regime we will discuss two cases: that of collisions and that of frequent collisions.

Ref. 2, the transition between these two cases is in $\omega = v_i/\epsilon$. The methods for solving the kinetic equation and for calculating Λ are described in detail in Ref. 2 (see also Ref. 7). We will reproduce only the results here.

The quantity Λ is given by

$$\Lambda = -i \frac{\sqrt{2} q^2 M_i}{\epsilon^2 n_0 \omega T} \int_0^{\infty} FE^{\frac{1}{2}} \left(1 - \frac{\hat{\omega}}{\omega}\right) dE \int \frac{d\lambda}{\lambda} \int d\theta \{[(1-\lambda B)^{\frac{1}{2}} + Y] - i\omega[(1-\lambda B)^{\frac{1}{2}} + Y]\}. \quad (19)$$

Here $\lambda = \mu/E$, $\mu = v_{\perp}^2/2B$, v_{\perp} is the modulus of the transverse velocity of the particles, and B is the equilibrium magnetic field. The integration over λ is from 0 to $1/B$; the region $0 < \lambda < 1/B_{\max}$ corresponds to untrapped particles, and the region $1/B_{\max} < \lambda < 1/B_{\min}$ to trapped particles. The integration over θ is from 0 to 2π in the case of untrapped particles and between the turning points in the case of the trapped particles. The function $Y = Y(\lambda, E)$ is nonzero only for the untrapped particles. It satisfies the relation

$$\int_0^{\infty} \frac{d\theta}{(1-\lambda B)^{\frac{1}{2}}} \{St[(1-\lambda B)^{\frac{1}{2}} + Y] + i\omega[(1-\lambda B)^{\frac{1}{2}} + Y]\} = 0. \quad (20)$$

The collision operator is

$$St(f) = 2\bar{v}_i \left(\frac{T}{M_i E}\right)^{\frac{1}{2}} \frac{(1-\lambda B)^{\frac{1}{2}}}{B} \frac{\partial}{\partial \lambda} \left[\lambda (1-\lambda B)^{\frac{1}{2}} \frac{\partial f}{\partial \lambda} \right], \quad (21)$$

where

$$\bar{v}_i = 3\sqrt{2} \pi v_i H(z)/2^{\frac{1}{2}}, \quad z = (M_i E/T)^{\frac{1}{2}},$$

$$H(z) = \frac{1}{\sqrt{\pi} z} \exp(-z^2) + \left(1 - \frac{1}{2z^2}\right) \frac{2}{\sqrt{\pi} z} \int_0^z \exp(-t^2) dt. \quad (22)$$

3.2.1. The case of infrequent collisions, $\omega > v_i/\epsilon$

Ignoring the collision term in Eq. (20), we find the following expression for the function Y :

$$Y = Y_0 = -\frac{\pi}{\sqrt{2}} \frac{\kappa \epsilon^{\frac{1}{2}}}{K(1/\kappa)}, \quad (23)$$

where $\kappa = (2\epsilon)^{-\frac{1}{2}} (1/\lambda B - 1 - \epsilon)^{\frac{1}{2}}$. In this case Λ becomes

$$\Lambda = \tau(1 - \omega_{pi}^*/\omega), \quad (24)$$

where

$$\tau = \frac{2^{\frac{1}{2}} q^2}{3\pi \epsilon^{\frac{1}{2}}} \left\{ 1 + \frac{9}{2} \int_0^{\infty} \kappa^2 d\kappa \left[E \left(\frac{1}{\kappa} \right) - \frac{\pi^2}{4K(1/\kappa)} \right] \right\}. \quad (25)$$

Here $K(1/\kappa)$ and $E(1/\kappa)$ are complete elliptic integrals of the first and second kinds, respectively. Now assuming that v_i is small but nonzero, we can incorporate collisions in Eq. (20). For this purpose we use the method of Ref. 8. We write the function Y in the form $Y = Y_0(1 + y)$; from Eq. (22) we then find

$$\frac{d^2 y}{dx^2} + i \frac{\epsilon \omega_0}{2\bar{v}_i} \left(\frac{EM_i}{T} \right)^{\frac{1}{2}} \ln \left(\frac{16}{|x|} \right) y = 0. \quad (26)$$

Here ω_0 is the real part of the frequency, and the parameter x is related to λ by $\lambda = [1 + \epsilon(2x - 1)]/\epsilon$

B. We seek the function y by making use of boundary conditions $y = -1$ at $x = 0$ and $y \rightarrow x \rightarrow -\infty$. From Eq. (26) we then find, in the quasiclassical approximation,

$$y = -\exp \left[\int_0^x \sigma(x') dx' \right],$$

where

$$\sigma(x) = \frac{1-i}{\sqrt{2}} \left[\frac{\epsilon \omega_0}{2\bar{v}_i} \left(\frac{M_i E}{T} \right)^{\frac{1}{2}} \ln \left(\frac{16}{|x|} \right) \right]^{\frac{1}{2}}.$$

Using expressions (19), (23), (27) and (28), find

$$\Lambda = \Lambda_0 + \Lambda_1,$$

where

$$\Lambda_1 = -(1+i) \mu \left(\frac{v_i}{\epsilon \omega_0} \right)^{\frac{1}{2}} \left[1 - \frac{\omega_{pi}^*}{\omega} + \eta \frac{\omega_{ni}^*}{\omega} \left(\frac{5}{2} - \frac{I_2}{I_1} \right) \right].$$

Here

$$\mu = \frac{\sqrt{3} \pi^{\frac{1}{2}} q^2 I_1}{2^{\frac{1}{2}} \epsilon^{\frac{1}{2}} \ln^{\frac{1}{2}}(128 \epsilon \omega_0 / v_i)},$$

and the numbers I_1 , and I_2 are defined by

$$I_1 = \int_0^{\infty} H^{\frac{1}{2}}(z) z^{\frac{1}{2}} \exp(-z^2) dz^2, \quad I_2 = \int_0^{\infty} H^{\frac{1}{2}}(z) z^{\frac{3}{2}} \exp(-z^2) dz^2.$$

An evaluation of these integrals yields $I_1 = 1$ and $I_2 = 0.74$.

Assuming $\gamma_0 \ll \omega_{pi}^*$, we find from (9), (30), and (31)

$$\omega = \omega_{pi}^* + 1.97i \left(\frac{v_i}{\epsilon \omega_{pi}^*} \right)^{\frac{1}{2}} \frac{\mu \eta}{1 + \tau} \omega_{ni}^*.$$

We see that an instability occurs in the case $\tau > 1.97 \left(\frac{v_i}{\epsilon \omega_{pi}^*} \right)^{\frac{1}{2}} \frac{\mu \eta}{1 + \tau} \omega_{ni}^*$. The ratio of the growth rate of the frequency $\sim (v_i/\epsilon \omega_{pi}^*)^{\frac{1}{2}}$.

Along with (33), the dispersion relation solution with $\omega < \omega_{pi}^*$. In this case we have

$$\omega = \omega_0 + i(v_i \omega_0 / \epsilon)^{\frac{1}{2}} \frac{\mu}{1 + \tau} \frac{1 - 0.97\eta}{1 + \eta},$$

where

$$\omega_0 = \gamma_0^2 / \omega_{pi}^* (1 + \tau).$$

An instability occurs if $-1 < \eta < 1.03$. The ratio of the growth rate to the frequency is $\sim (v_i/\epsilon \omega_{pi}^*)^{\frac{1}{2}}$.

3.2.2. The case of frequent collisions. Ignoring the term with ω in Eq. (20), we find

$$\frac{1}{B} \frac{\partial Y}{\partial \lambda} = \frac{\pi}{2^{\frac{1}{2}}} \frac{1}{\kappa E(1/\kappa) \epsilon^{\frac{1}{2}}}.$$

From (19) and (36) we then find

$$\Lambda = i \frac{q^2 \pi}{4\epsilon^{\frac{1}{2}}} \frac{3}{2} \frac{v_i}{\omega} \left\{ [\sqrt{2} - \ln(1 + \sqrt{2})] - \frac{\omega_{ni}^*}{\omega} \left[\left(1 - \frac{3}{2} \eta\right) \right] \right\}$$

$$-\ln(1+\sqrt{2}) + \frac{1}{\sqrt{2}}\eta \left\{ 1 - 2 \int_0^1 dx \left[\frac{1}{E(1/x)} - \frac{4}{\pi^2} K\left(\frac{1}{x}\right) \right] \right\}. \quad (37)$$

From (9) and (37) we find the wave branch

$$\omega = \omega_{ni} [1 + \eta(1-k)] - i2 \frac{e^{\eta} \omega_{ni}^4}{q^2 v_i} k \eta [1 + \eta(1-k)]. \quad (38)$$

Here $k = 1.03$. We see that this branch is unstable if $\eta < 0$.

4. DISCUSSION OF RESULTS

This analysis indicates the possibility of non-hydromagnetic instabilities of drift-ballooning modes as a result of resonant and collisional interactions of waves with ions. Correspondingly, these instabilities could be called "resonant" and "collisional." The resonant instabilities are characterized by (11), (15), and (16); the collisional instabilities by (33), (34), and (38). Both depend strongly on the parameter $\eta = \partial \ln T / \partial \ln n_0$ [see, e.g., conditions (12), (17) and (18)].

These instabilities may be of interest for the theory of plasma turbulence in tokamaks, in particular, because they may — by virtue of the non-linearity of the plasma equations — influence the longer-wavelength (hydromagnetically stable) perturbations. Furthermore, hydromagnetic stability can be achieved not only by virtue of the finite-Larmor-radius effects but also by imposing an ex-

ternal agent on the plasma, e.g., by injecting particles into the tokamak.⁹ Under these conditions the instabilities which we have been discussing may be the most important instabilities. These arguments will of course require a detailed theory, but that would go beyond the scope of the present paper.

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